# A non-existence result for supersonic travelling waves in the Gross-Pitaevskii equation 

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#### Abstract

We prove the non-existence of non-constant travelling waves of finite energy and of speed $c>\sqrt{2}$ in the Gross-Pitaevskii equation in dimension $N \geq 2$.


## Introduction

In this paper, we will focus on the Gross-Pitaevskii equation

$$
\begin{equation*}
i \partial_{t} u=\Delta u+u\left(1-|u|^{2}\right) . \tag{1}
\end{equation*}
$$

One of the motivations for this equation is the analysis of Bose-Einstein condensation, which describes the behaviour of interacting bosons near absolute zero. When condensation occurs, equation (1) might be used as a model for the Bose condensate (see [4] for more details). In particular, this model is relevant to describe Bose-condensed gases. The model is also sometimes proposed to describe the superfluid state of Helium II, though in this case the interactions between particles are important and cannot be neglected at temperature different from zero.

In order to describe this condensation, E.P. Gross [8] and L.P. Pitaevskii [12] considered a set of $N$ bosons of mass $m$ that fill a volume $V$ : they then assumed almost all bosons are Bosecondensed in the fundamental state of energy. Therefore, they can be described by a macroscopic wave function $\Psi$. They then deduced the Gross-Pitaevskii equation satisfied by the function $\Psi$ from a Hartree-Fock approach,

$$
i \hbar \partial_{t} \Psi+\frac{\hbar^{2}}{2 m} \Delta \Psi-\Psi \int_{V}\left|\Psi\left(x^{\prime}, t\right)\right|^{2} U\left(x-x^{\prime}\right) d x^{\prime}=0 .
$$

Here, $U\left(x-x^{\prime}\right)$ denotes the interaction between the bosons at positions $x$ and $x^{\prime}$ : this interaction being of very short range, it is often approached by $U_{0} \delta\left(x-x^{\prime}\right)$. Thus, denoting $E_{b}$, the average energy level per unit mass of a boson, and,

$$
u(t, x)=e^{\frac{-i m E_{b} t}{\hbar}} \Psi(t, x),
$$

they computed the equation

$$
i \hbar \partial_{t} u+\frac{\hbar^{2}}{2 m} \Delta u+m E_{b} u-U_{0} u|u|^{2}=0 .
$$

They finally rescaled the equation by taking the mean density $\rho_{0}=\sqrt{\frac{m E_{b}}{U_{0}}}$ as unity, $\frac{\hbar}{\sqrt{2 m^{2} E_{b}}}$ as unit length, and, $\frac{\hbar}{m E_{b}}$ as unit time, in order to obtain the dimensionless equation

$$
\begin{equation*}
i \partial_{t} u+\Delta u+u\left(1-|u|^{2}\right)=0 . \tag{1}
\end{equation*}
$$

[^0]At this point, we can write the hydrodynamic form of this equation by using the Madelung transform [11],

$$
u=\sqrt{\rho} e^{i \theta},
$$

which is only meaningful where $\rho$ does not vanish. Denoting $v=2 \nabla \theta$, we deduce the equations

$$
\left\{\begin{array}{l}
\partial_{t} \rho+\operatorname{div}(\rho v)=0, \\
\rho\left(\partial_{t} v+v \cdot \nabla v\right)+\nabla \rho^{2}=\rho \nabla\left(\frac{\Delta \rho}{\rho}-\frac{|\nabla \rho|^{2}}{2 \rho^{2}}\right) .
\end{array}\right.
$$

Those equations are similar to the Euler equations for a irrotational ideal fluid with pressure $p(\rho)=\rho^{2}$ : the term of the right member is then considered as a quantic pressure term. Here, we can remark that the sound speed is $c_{s}=\sqrt{2}$.

In this article, we will consider equation (1) in the space $\mathbb{R}^{N}$ for every integer $N \geq 2$ : we can notice that this equation is associated to the energy

$$
E(u)=\frac{1}{2} \int_{\mathbb{R}^{N}}|\nabla u|^{2}+\frac{1}{4} \int_{\mathbb{R}^{N}}\left(1-|u|^{2}\right)^{2}=\int_{\mathbb{R}^{N}} e(u) .
$$

We will study the travelling waves of finite energy and of speed $c \geq 0$ for this equation i.e. the solutions $u$ which are of the form

$$
u(t, x)=v\left(x_{1}-c t, \ldots, x_{N}\right)
$$

The simplified equation for $v$, which we will consider now, is

$$
\begin{equation*}
i c \partial_{1} v+\Delta v+v\left(1-|v|^{2}\right)=0 \tag{2}
\end{equation*}
$$

C.A. Jones, S.J. Putterman and P.H. Roberts [10, 9] first considered formally and numerically those particular solutions because they suppose they play an important role in the long time dynamics of general solutions: they conjectured that non-constant travelling waves only exist when their speed $c$ is in the interval $] 0, \sqrt{2}[$ ie they all are subsonic. They then noticed the apparition of vortices for those solutions when $c$ tends to 0 in dimension two (two parallel oppositely directed vortices) and in dimension three (a vortex ring). They also gave for each value of $c$, the asymptotic development at infinity in dimension two,

$$
v(x)-1 \underset{|x| \rightarrow+\infty}{\sim} \frac{i \alpha x_{1}}{x_{1}^{2}+\left(1-\frac{c^{2}}{2}\right) x_{2}^{2}}
$$

and in dimension three,

$$
v(x)-1 \underset{|x| \rightarrow+\infty}{\sim} \frac{i \alpha x_{1}}{\left(x_{1}^{2}+\left(1-\frac{c^{2}}{2}\right)\left(x_{2}^{2}+x_{3}^{2}\right)\right)^{\frac{3}{2}}},
$$

where the constant $\alpha$ is the stretched dipole coefficient.
F. Béthuel and J.C. Saut [3, 2] first studied mathematically those travelling waves: they proved their existence in dimension two when $c$ is small and the apparition of vortices in this case. They also gave a mathematical proof for their limit at infinity.
In dimension $N \geq 3$, F. Béthuel, G. Orlandi and D. Smets [1] showed their existence when $c$ is small and the apparition of a vortex ring.
In every dimension, A. Farina [5] proved a universal bound for their modulus.
Finally, we proved their uniform convergence to a constant of modulus one in dimension $N \geq 3$ [6], and also studied their decay at infinity in dimension $N \geq 2$ [7].

In this paper, we will complete all those results by the following theorem.

Theorem 1. In dimension $N \geq 2$, a solution of equation (2) of finite energy and speed $c>\sqrt{2}$ is constant.

This paper will be organised around the proof of Theorem 1. In the first step, we will write the equation satisfied by $\eta=1-|v|^{2}$. Then, we will derive a new integral identity when $c>\sqrt{2}$ : this is the crucial step of the proof of Theorem 1. Finally, we will write the Pohozaev identities in order to prove that the energy $E(v)$ vanishes and that the travelling wave $v$ is constant.

## 1 Equation satisfied by $\eta$

In this part, we will write the equation satisfied by the variable $\eta=1-|v|^{2}$ for every $c \geq 0$ : in particular, the results in this section (i.e. Propositions 1,2 and 3 ) are valid for every $c \geq 0$. We first recall two useful propositions yet mentioned in $[6,7]$ and based on arguments taken from F. Béthuel and J.C. Saut [3, 2].

Proposition 1. For every $c \geq 0$, if $v$ is a solution of equation (2) in $L_{l o c}^{1}\left(\mathbb{R}^{N}\right)$ of finite energy, then $v$ is regular, bounded and its gradient belongs to all the spaces $W^{k, p}\left(\mathbb{R}^{N}\right)$ for $k \in \mathbb{N}$ and $p \in[2,+\infty]$.

Thus, a travelling wave is a regular function and a classical solution of equation (2), which will simplify the following discussion.

Proposition 2. The modulus $\rho$ of $v$ satisfies

$$
\rho(x) \underset{|x| \rightarrow+\infty}{\rightarrow} 1 .
$$

Proof. Indeed, the function $\eta^{2}$ is uniformly continuous because $v$ is bounded and lipschitzian by Proposition 1. As $\int_{\mathbb{R}^{N}} \eta^{2}$ is finite, $\eta$ converges uniformly to 0 at infinity, which completes the proof of this proposition.

Thus, the function $\rho$ does not vanish at infinity, and we can define a regular function $\theta$ on a neighborhood of infinity such that $v$ can be written

$$
v=\rho e^{i \theta} .
$$

Denoting $\psi$, a regular function from $\mathbb{R}^{N}$ to $[0,1]$ such that $\psi=0$ on a neighborhood of $Z=$ $\left\{x \in \mathbb{R}^{N}, \rho(x)=0\right\}$, and $\psi=1$ on a neighbourhood of infinity, and denoting $v=v_{1}+i v_{2}$, we can write the equation satisfied by the function $\eta$ :

Proposition 3. For every $c \geq 0$, the function $\eta$ satisfies the equation

$$
\begin{equation*}
\Delta^{2} \eta-2 \Delta \eta+c^{2} \partial_{1,1}^{2} \eta=-\Delta F+2 c \partial_{1} \operatorname{div}(\mathrm{G}) \tag{3}
\end{equation*}
$$

where

$$
F=2|\nabla v|^{2}+2 \eta^{2}+2 c\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}\right)-2 c \partial_{1}(\psi \theta)
$$

and

$$
G=v_{1} \nabla v_{2}-v_{2} \nabla v_{1}-\nabla(\psi \theta) .
$$

Proof. By equation (2), we have

$$
\begin{align*}
& \Delta v_{1}-c \partial_{1} v_{2}+v_{1}\left(1-|v|^{2}\right)=0  \tag{4}\\
& \Delta v_{2}+c \partial_{1} v_{1}+v_{2}\left(1-|v|^{2}\right)=0 \tag{5}
\end{align*}
$$

We then compute

$$
\Delta^{2} \eta-2 \Delta \eta+c^{2} \partial_{1,1}^{2} \eta=-2 \Delta|\nabla v|^{2}-2 \Delta(v . \Delta v)-2 \Delta \eta+c^{2} \partial_{1,1}^{2} \eta
$$

and by equations (4)-(5), we have on one hand

$$
v . \Delta v=v_{1} \Delta v_{1}+v_{2} \Delta v_{2}=c\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}\right)-|v|^{2} \eta
$$

and, on the other hand,

$$
c \partial_{1} \eta=-2 c\left(v_{1} \partial_{1} v_{1}+v_{2} \partial_{1} v_{2}\right)=2\left(\Delta v_{2} v_{1}-\Delta v_{1} v_{2}\right)=2 \operatorname{div}\left(\nabla v_{2} v_{1}-\nabla v_{1} v_{2}\right)
$$

Therefore, we finally get

$$
\begin{aligned}
\Delta^{2} \eta-2 \Delta \eta+c^{2} \partial_{1,1}^{2} \eta & =-2 \Delta|\nabla v|^{2}-2 \Delta \eta^{2}-2 c \Delta\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}\right) \\
& +2 c \partial_{1} \operatorname{div}\left(v_{1} \nabla v_{2}-v_{2} \nabla v_{1}\right) \\
& =-\Delta\left(2|\nabla v|^{2}+2 \eta^{2}+2 c\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}\right)-2 c \partial_{1}(\psi \theta)\right) \\
& +2 c \partial_{1} \operatorname{div}\left(v_{1} \nabla v_{2}-v_{2} \nabla v_{1}-\nabla(\psi \theta)\right) \\
& =-\Delta F+2 c \partial_{1} \operatorname{div}(G)
\end{aligned}
$$

which is the desired equality.

## 2 A new integral relation

We have
Proposition 4. If $c>\sqrt{2}$, the travelling wave $v$ satisfies the integral equation

$$
\begin{equation*}
\int_{\mathbb{R}^{N}}\left(|\nabla v|^{2}+\eta^{2}\right)=c\left(\frac{2}{c^{2}}-1\right) \int_{\mathbb{R}^{N}}\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}-\partial_{1}(\psi \theta)\right) \tag{6}
\end{equation*}
$$

Remark 1. This is the only point where we use the assumption $c>\sqrt{2}$.
For the proof, we use
Lemma 1. $F$ and $G$ belong to the space $W^{2,1}\left(\mathbb{R}^{N}\right)$.
Proof. Indeed, $G$ is regular and satisfies at infinity

$$
G=\left(\rho^{2}-1\right) \nabla \theta
$$

By Proposition 1 , the functions $\eta$ and $\nabla v$ belong to $H^{2}\left(\mathbb{R}^{N}\right) \cap W^{2, \infty}\left(\mathbb{R}^{N}\right)$ : since

$$
|\nabla v|^{2}=|\nabla \rho|^{2}+\rho^{2}|\nabla \theta|^{2}
$$

and since $\rho$ uniformly converges to 1 at infinity by Proposition 2 , the function $\nabla \theta$ belongs to $H^{2} \cap W^{2, \infty}$ on a neighbourhood of infinity: thus, the function $G$ belongs to the space $W^{2,1}\left(\mathbb{R}^{N}\right)$. Since

$$
F=2\left(|\nabla v|^{2}+\eta^{2}\right)+2 c G_{1}
$$

the function $F$ also belongs to this space, which completes the proof of Lemma 1.

Proof of Proposition 4. By Proposition 1, the function $\eta$ belongs to $H^{4}\left(\mathbb{R}^{N}\right)$, and we can write by taking the Fourier transformation of equation (3)

$$
\begin{equation*}
\forall \xi \in \mathbb{R}^{N},\left(|\xi|^{4}+2|\xi|^{2}-c^{2} \xi_{1}^{2}\right) \widehat{\eta}(\xi)=|\xi|^{2} \widehat{F}(\xi)-2 c \sum_{j=1}^{N} \xi_{1} \xi_{j} \widehat{G_{j}}(\xi):=H(\xi) \tag{7}
\end{equation*}
$$

Consider the set

$$
\Gamma=\left\{\xi \in \mathbb{R}^{N},|\xi|^{4}+2|\xi|^{2}-c^{2} \xi_{1}^{2}=0\right\}
$$

This set is reduced to $\{0\}$ when $c \leq \sqrt{2}$, but, when $c>\sqrt{2}$, it is a regular hypersurface of codimension 1 except at $\{0\}$ : in dimension 2 , it has the geometry of a bretzel, and in higher dimensions, it has the geometry of two spheres linked at some point. Indeed, $\Gamma$ is a surface of revolution around axis $x_{1}$ : in spherical coordinates $\xi=(r \cos (\alpha), r \sin (\alpha) \cos (\beta), \ldots)$, it is described by the equation

$$
r^{2}=c^{2} \cos ^{2}(\alpha)-2
$$

In particular, we notice that there are two sequences $\left(x_{n}\right)_{n \in \mathbb{N}}$ and $\left(y_{n}\right)_{n \in \mathbb{N}}$ of points of $\Gamma \backslash\{0\}$ which tend to 0 when $n$ tends to $+\infty$ and which satisfy

$$
\begin{equation*}
\frac{x_{n}}{\left|x_{n}\right|} \underset{n \rightarrow+\infty}{\rightarrow}\left(\frac{2}{c^{2}}, \sqrt{1-\frac{2}{c^{2}}}, 0, \ldots\right), \text { and }, \frac{y_{n}}{\left|y_{n}\right|} \underset{n \rightarrow+\infty}{\rightarrow}\left(\frac{2}{c^{2}},-\sqrt{1-\frac{2}{c^{2}}}, 0, \ldots\right) \tag{8}
\end{equation*}
$$

Coming back to the study of equation (7), we claim that
Lemma 2. The function $H$ defined by equation (7) is continuous on $\mathbb{R}^{N}$ and satisfies

$$
H=0 \text { on } \Gamma
$$

Proof of Lemma 2. The first assertion follows from Lemma 1: indeed, since the functions $F$ and $G$ belong to the space $W^{2,1}\left(\mathbb{R}^{N}\right)$, the functions $\xi \mapsto|\xi|^{2} \widehat{F}(\xi)$ and $\xi \mapsto \xi_{1} \xi_{j} \widehat{G_{j}}(\xi)$ are continuous on $\mathbb{R}^{N}$, and therefore, the function $H$ is continuous on $\mathbb{R}^{N}$ too.
In order to prove the second assertion, we argue by contradiction and assume there is some point $\xi_{0} \in \Gamma \backslash\left\{0,\left(\sqrt{c^{2}-2}, 0, \ldots, 0\right)\right\}$ such that

$$
H\left(\xi_{0}\right) \neq 0
$$

Since the function $H$ is continuous on $\mathbb{R}^{N}$, there is some neighbourhood $V$ of the point $\xi_{0}$ and some strictly positive number $A$ such that

$$
\forall \xi \in V,|H(\xi)| \geq A
$$

Hence, we have by equation (7)

$$
\forall \xi \in V \backslash \Gamma,|\widehat{\eta}(\xi)|^{2} \geq \frac{A^{2}}{\left(|\xi|^{4}+2|\xi|^{2}-c^{2} \xi_{1}^{2}\right)^{2}}
$$

Integrating this relation and using spherical coordinates, we get

$$
\begin{aligned}
\int_{V \backslash \Gamma}|\widehat{\eta}(\xi)|^{2} d \xi & \geq A^{2} \int_{V \backslash \Gamma} \frac{d \xi}{\left(|\xi|^{4}+2|\xi|^{2}-c^{2} \xi_{1}^{2}\right)^{2}} \\
& \geq A_{N} \int_{V \backslash \Gamma \cap \mathbb{R} \times \mathbb{R}_{+}} \frac{s^{N-2} d s d \xi_{1}}{\left(\left(s^{2}+\xi_{1}^{2}\right)^{2}+2 s^{2}+\left(2-c^{2}\right) \xi_{1}^{2}\right)^{2}} \\
& \geq A_{N} \int_{V \backslash \Gamma \cap \mathbb{R}_{+} \times[0, \pi]} \frac{r^{N-1} \sin ^{N-2}(\alpha) d r d \alpha}{r^{4}\left(r^{2}+2-c^{2} \cos ^{2}(\alpha)\right)^{2}}
\end{aligned}
$$

Thus, denoting

$$
\xi_{0}=\left(r_{0} \cos \left(\alpha_{0}\right), r_{0} \sin \left(\alpha_{0}\right) \cos \left(\beta_{0}\right), \ldots\right),
$$

there is some real number $\epsilon>0$ such that

$$
\int_{V \backslash \Gamma}|\widehat{\eta}(\xi)|^{2} d \xi \geq A_{N} \int_{r_{0}-\epsilon}^{r_{0}+\epsilon} \int_{\alpha_{0}-\epsilon}^{\alpha_{0}+\epsilon} \frac{r^{N-1} \sin ^{N-2}(\alpha) d r d \alpha}{r^{4}\left(r^{2}+2-c^{2} \cos ^{2}(\alpha)\right)^{2}}:=A_{N} I\left(\alpha_{0}, r_{0}, \epsilon\right) .
$$

Since $\xi_{0} \in \Gamma \backslash\left\{0,\left(\sqrt{c^{2}-2}, 0, \ldots, 0\right)\right\}, r_{0}$ is different from 0 and $\alpha_{0}$ is different from 0 and $\frac{\pi}{2}$, and so, we can compute for $\epsilon$ sufficiently small

$$
I\left(\alpha_{0}, r_{0}, \epsilon\right) \geq A\left(\alpha_{0}, r_{0}, \epsilon\right) \int_{r_{0}-\epsilon}^{r_{0}+\epsilon} \int_{\alpha_{0}-\epsilon}^{\alpha_{0}+\epsilon} \frac{d r d \alpha}{\left(r^{2}+2-c^{2} \cos ^{2}(\alpha)\right)^{2}} .
$$

By doing the change of variable $r=\sqrt{c^{2} \cos ^{2}(\beta)-2}$, we know that there is some real number $\delta>0$ such that

$$
I\left(\alpha_{0}, r_{0}, \epsilon\right) \geq A\left(\alpha_{0}, r_{0}, \epsilon\right) \int_{\alpha_{0}-\delta}^{\alpha_{0}+\delta} \int_{\alpha_{0}-\delta}^{\alpha_{0}+\delta} \frac{d \beta d \alpha}{\left(c^{2} \cos ^{2}(\beta)-c^{2} \cos ^{2}(\alpha)\right)^{2}},
$$

and finally, by denoting $a=\alpha-\alpha_{0}$ and $b=\beta-\alpha_{0}$, we get

$$
\begin{aligned}
I\left(\alpha_{0}, r_{0}, \epsilon\right) & \geq A\left(\alpha_{0}, r_{0}, \epsilon, c\right) \int_{-\delta}^{\delta} \int_{-\delta}^{\delta} \frac{d a d b}{\left(\cos ^{2}\left(b+\alpha_{0}\right)-\cos ^{2}\left(a+\alpha_{0}\right)\right)^{2}} \\
& \geq A\left(\alpha_{0}, r_{0}, \epsilon, c\right) \int_{-\delta}^{\delta} \int_{-\delta}^{\delta} \frac{d a d b}{\left(\cos \left(2 b+2 \alpha_{0}\right)-\cos \left(2 a+2 \alpha_{0}\right)\right)^{2}} \\
& \geq A\left(\alpha_{0}, r_{0}, \epsilon, c\right) \int_{-\delta}^{\delta} \int_{-\delta}^{\delta} \frac{d a d b}{(\sin (b-a))^{2}} .
\end{aligned}
$$

Since the function $(a, b) \mapsto \frac{1}{(\sin (b-a))^{2}}$ is not integrable at the origin in $\mathbb{R}^{2}$, the integral $I\left(\alpha_{0}, r_{0}, \epsilon\right)$ is not finite and we can conclude that

$$
\int_{V \backslash \Gamma}|\widehat{\eta}(\xi)|^{2} d \xi=+\infty .
$$

Since the energy of the function $v$ is finite, so is the integral $\int_{\mathbb{R}^{N}} \eta^{2}$, and by Plancherel theorem, we deduce

$$
\int_{\mathbb{R}^{N}}|\widehat{\eta}(\xi)|^{2} d \xi<+\infty
$$

which leads to a contradiction and proves that $H$ is identically equal to 0 on the set $\Gamma \backslash$ $\left\{0,\left(\sqrt{c^{2}-2}, 0, \ldots, 0\right)\right\}$. The second assertion of Lemma 2 then follows from the continuity of the function $H$.

End of the proof of Proposition 4. By Lemma 2, we now know that

$$
\forall n \in \mathbb{N}, H\left(x_{n}\right)=0,
$$

which gives by dividing by $\left|x_{n}\right|^{2}$,

$$
\forall n \in \mathbb{N}, \widehat{F}\left(x_{n}\right)=2 c \sum_{j=1}^{N} \frac{\left(x_{n}\right)_{1}}{\left|x_{n}\right|} \frac{\left(x_{n}\right)_{j}}{\left|x_{n}\right|} \widehat{G_{j}}\left(x_{n}\right)
$$

By continuity of the functions $\widehat{F}$ and $\widehat{G_{j}}$, we can take the limit as $x_{n} \rightarrow 0$ of this expression and obtain by assertion 8

$$
\widehat{F}(0)=\frac{4}{c} \widehat{G_{1}}(0)+2 c \sqrt{1-\frac{2}{c^{2}}} \widehat{G_{2}}(0)
$$

Likewise, we know that

$$
\forall n \in \mathbb{N}, H\left(y_{n}\right)=0
$$

which gives by the same method,

$$
\widehat{F}(0)=\frac{4}{c} \widehat{G_{1}}(0)-2 c \sqrt{1-\frac{2}{c^{2}}} \widehat{G_{2}}(0)
$$

Finally, we have

$$
\widehat{F}(0)=\frac{4}{c} \widehat{G_{1}}(0)
$$

so that,

$$
\int_{\mathbb{R}^{N}} F(x) d x=\frac{4}{c} \int_{\mathbb{R}^{N}} G_{1}(x) d x
$$

The conclusion follows from the expression of the functions $F$ and $G$.

## 3 Pohozaev identities

We now prove for sake of completeness two well-known identities based on the use of Pohozaev multipliers (See $[10,9,3,1]$ for more details). Those estimates do not use the fact that $c>\sqrt{2}$.

Proposition 5. Let $c \geq 0$. A finite energy solution $v$ to equation (2) satisfies the two identities

$$
\begin{align*}
& E(v)=\int_{\mathbb{R}^{N}}\left|\partial_{1} v\right|^{2}  \tag{9}\\
& \forall 2 \leq j \leq N, E(v)=\int_{\mathbb{R}^{N}}\left|\partial_{j} v\right|^{2}+\frac{c}{2} \int_{\mathbb{R}^{N}}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right) \tag{10}
\end{align*}
$$

Proof. We first fix some real number $R>0$ and we multiply equation (2) by Pohozaev multiplier $x_{1} \partial_{1} v$ on the ball $B(0, R)$

$$
\begin{equation*}
\int_{B(0, R)}\left(\Delta v \cdot x_{1} \partial_{1} v+x_{1} \partial_{1} v \cdot v\left(1-|v|^{2}\right)\right)=0 \tag{11}
\end{equation*}
$$

Integrating by parts, we compute

$$
\begin{aligned}
\int_{B(0, R)} \Delta v \cdot x_{1} \partial_{1} v & =\int_{B(0, R)} \frac{|\nabla v|^{2}}{2}-\int_{B(0, R)}\left|\partial_{1} v\right|^{2}+\int_{S(0, R)} x_{1} \partial_{1} v \cdot \partial_{\nu} v \\
& -\int_{S(0, R)} \nu_{1} x_{1} \frac{|\nabla v|^{2}}{2}
\end{aligned}
$$

and

$$
\int_{B(0, R)} x_{1} \partial_{1} v \cdot v\left(1-|v|^{2}\right)=\int_{B(0, R)} \frac{\left(1-|v|^{2}\right)^{2}}{4}-\int_{S(0, R)} x_{1} \nu_{1} \frac{\left(1-|v|^{2}\right)^{2}}{4}
$$

By equation (11), we then get

$$
\begin{equation*}
\int_{B(0, R)} e(v)=\int_{B(0, R)}\left|\partial_{1} v\right|^{2}-\int_{S(0, R)} x_{1} \partial_{1} v \cdot \partial_{\nu} v+\int_{S(0, R)} \nu_{1} x_{1} e(v) \tag{12}
\end{equation*}
$$

On one hand, by Proposition 1, we know that

$$
\int_{B(0, R)} e(v)-\int_{B(0, R)}\left|\partial_{1} v\right|^{2} \underset{R \rightarrow+\infty}{\rightarrow} E(v)-\int_{\mathbb{R}^{N}}\left|\partial_{1} v\right|^{2} .
$$

On the other hand, we have

$$
\left|\int_{S(0, R)} x_{1} \partial_{1} v \cdot \partial_{\nu} v-\nu_{1} x_{1} e(v)\right| \leq A R \int_{S(0, R)} e(v) .
$$

Since the integral $\int_{\mathbb{R}_{+}}\left(\int_{S(0, R)} e(v)\right) d R$ is finite, there are some positive real numbers $R_{n}$ such that $R_{n} \xrightarrow[n \rightarrow+\infty]{ }+\infty$ and

$$
\forall n \in \mathbb{N}, R_{n} \int_{S\left(0, R_{n}\right)} e(v) \leq \frac{1}{\ln \left(R_{n}\right)},
$$

which gives

$$
\int_{S\left(0, R_{n}\right)}\left(x_{1} \partial_{1} v \cdot \partial_{\nu} v-\nu_{1} x_{1} e(v)\right) \underset{n \rightarrow+\infty}{\rightarrow} 0
$$

and finally, by equation (12),

$$
E(v)=\int_{\mathbb{R}^{N}}\left|\partial_{1} v\right|^{2} .
$$

In order to prove the second identity, we multiply equation (2) by Pohozaev multiplier $x_{j} \partial_{j} v$ on the ball $B(0, R)$

$$
\begin{equation*}
\int_{B(0, R)}\left(\Delta v \cdot x_{j} \partial_{j} v+i c \partial_{1} v \cdot x_{j} \partial_{j} v+x_{j} \partial_{j} v \cdot v\left(1-|v|^{2}\right)\right)=0 . \tag{13}
\end{equation*}
$$

Integrating by parts, we compute

$$
\begin{aligned}
\int_{B(0, R)} \Delta v \cdot x_{j} \partial_{j} v & =\int_{B(0, R)} \frac{|\nabla v|^{2}}{2}-\int_{B(0, R)}\left|\partial_{j} v\right|^{2}+\int_{S(0, R)} x_{j} \partial_{j} v \cdot \partial_{\nu} v \\
& -\int_{S(0, R)} \nu_{j} x_{j} \frac{|\nabla v|^{2}}{2}
\end{aligned}
$$

and

$$
\int_{B(0, R)} x_{j} \partial_{j} v \cdot v\left(1-|v|^{2}\right)=\int_{B(0, R)} \frac{\left(1-|v|^{2}\right)^{2}}{4}-\int_{S(0, R)} x_{j} \nu_{j} \frac{\left(1-|v|^{2}\right)^{2}}{4} .
$$

If $R$ is sufficiently large such as $\psi=1$ on $S(0, R)$, we also compute

$$
\begin{aligned}
\int_{B(0, R)} i \partial_{1} v \cdot x_{j} \partial_{j} v & =\frac{1}{2}\left(\int_{S(0, R)} x_{j}\left(\rho^{2}-1\right)\left(\nu_{j} \partial_{1} \theta-\nu_{1} \partial_{j} \theta\right)\right. \\
& \left.-\int_{B(0, R)}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right)\right),
\end{aligned}
$$

which leads to

$$
\begin{align*}
\int_{B(0, R)} e(v) & =\int_{B(0, R)}\left|\partial_{j} v\right|^{2}+\frac{c}{2} \int_{B(0, R)}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right) \\
& +\int_{S(0, R)}\left(x_{j} \nu_{j} e(v)-x_{j} \partial_{j} v \cdot \partial_{\nu} v-x_{j}\left(\rho^{2}-1\right)\left(\nu_{j} \partial_{1} \theta-\nu_{1} \partial_{j} \theta\right)\right) . \tag{14}
\end{align*}
$$

On one hand, by Proposition 1, we know that

$$
\begin{aligned}
& \int_{B(0, R)} e(v)-\int_{B(0, R)}\left|\partial_{j} v\right|^{2}-\frac{c}{2} \int_{B(0, R)}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right) \\
& \underset{R \rightarrow+\infty}{\rightarrow} E(v)-\int_{\mathbb{R}^{N}}\left|\partial_{j} v\right|^{2}-\frac{c}{2} \int_{\mathbb{R}^{N}}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right)
\end{aligned}
$$

On the other hand, we have

$$
\left|\int_{S(0, R)}\left(x_{j} \nu_{j} e(v)-x_{j} \partial_{j} v \cdot \partial_{\nu} v-x_{j}\left(\rho^{2}-1\right)\left(\nu_{j} \partial_{1} \theta-\nu_{1} \partial_{j} \theta\right)\right)\right| \leq A R \int_{S(0, R)} e(v)
$$

By using the sequence of positive real numbers $R_{n}$ constructed for proving equality (9), we get

$$
\int_{S\left(0, R_{n}\right)}\left(x_{j} \nu_{j} e(v)-x_{j} \partial_{j} v . \partial_{\nu} v-x_{j}\left(\rho^{2}-1\right)\left(\nu_{j} \partial_{1} \theta-\nu_{1} \partial_{j} \theta\right)\right) \underset{n \rightarrow+\infty}{\rightarrow} 0
$$

and finally, by equation (14),

$$
E(v)=\int_{\mathbb{R}^{N}}\left|\partial_{j} v\right|^{2}+\frac{c}{2} \int_{\mathbb{R}^{N}}\left(v_{2} \partial_{1} v_{1}-v_{1} \partial_{1} v_{2}+\partial_{1}(\psi \theta)\right)
$$

## 4 Conclusion

We now complete the proof of Theorem 1. By Proposition 5, we have

$$
E(v)=\int_{\mathbb{R}^{N}}\left|\partial_{1} v\right|^{2}
$$

which gives by denoting $\nabla_{\perp} v=\left(\partial_{2} v, \ldots, \partial_{N} v\right)$,

$$
\frac{1}{2} \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2}+\frac{1}{4} \int_{\mathbb{R}^{N}} \eta^{2}=\frac{1}{2} \int_{\mathbb{R}^{N}}\left|\partial_{1} v\right|^{2}=\frac{E(v)}{2}
$$

and

$$
\begin{equation*}
\int_{\mathbb{R}^{N}} \eta^{2}=2 E(v)-2 \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2} \tag{15}
\end{equation*}
$$

We then compute

$$
\begin{equation*}
\int_{\mathbb{R}^{N}}\left(|\nabla v|^{2}+\eta^{2}\right)=3 E(v)-\int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2} \tag{16}
\end{equation*}
$$

and, by Proposition 5,

$$
\begin{equation*}
c \int_{\mathbb{R}^{N}}\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}-\partial_{1}(\psi \theta)\right)=\frac{2}{N-1} \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2}-2 E(v) \tag{17}
\end{equation*}
$$

Proposition 4 gives

$$
\int_{\mathbb{R}^{N}}\left(|\nabla v|^{2}+\eta^{2}\right)=c\left(\frac{2}{c^{2}}-1\right) \int_{\mathbb{R}^{N}}\left(v_{1} \partial_{1} v_{2}-v_{2} \partial_{1} v_{1}-\partial_{1}(\psi \theta)\right)
$$

which leads by equations (16)-(17) to

$$
\begin{equation*}
\left(c^{2}+4\right)(N-1) E(v)=\left((N-3) c^{2}+4\right) \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2} \tag{18}
\end{equation*}
$$

If $N=2$, we get

$$
\left(c^{2}+4\right) E(v)=\left(4-c^{2}\right) \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2},
$$

which gives by equation (15),

$$
\frac{c^{2}+4}{2} \int_{\mathbb{R}^{N}} \eta^{2}=-2 c^{2} \int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2}=0 .
$$

Finally, we have involving equation (15) once more

$$
E(v)=0 .
$$

If $N \geq 3$, since by equation (15),

$$
\int_{\mathbb{R}^{N}}\left|\nabla_{\perp} v\right|^{2} \leq E(v),
$$

equation (18) gives

$$
\left(2 c^{2}+4 N-8\right) E(v) \leq 0,
$$

and finally, $E(v)$ is also equal to 0 in this case.
In conclusion, since $E(v)=0$, the function $\nabla v$ vanishes on $\mathbb{R}^{N}$ and $v$ is a constant (of modulus one since $\eta$ also vanishes on $\mathbb{R}^{N}$ ).

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